Our Solar System Learning Guide

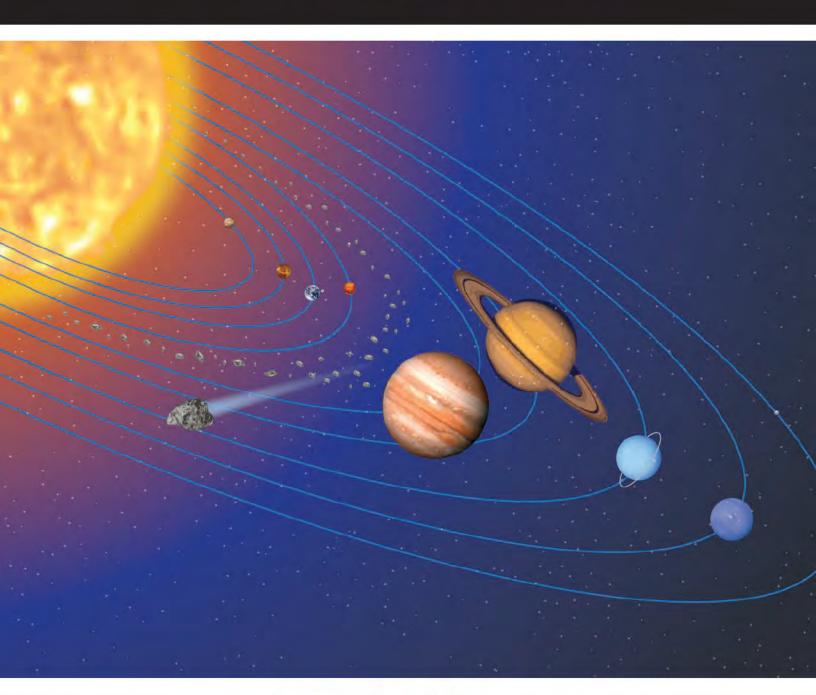




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FORMATION OF OUR SOLAR SYSTEM

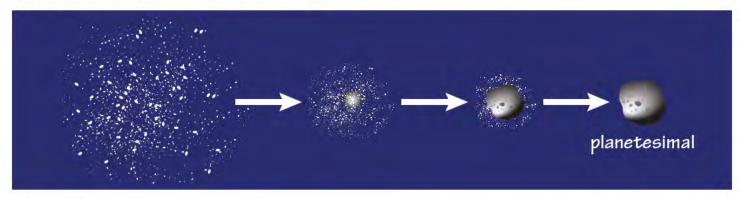
Nebulae

Solar systems begin in dusty gas clouds called **nebulae** that are found between stars. The dust is composed of elements like **iron** and **carbon**. Nebula gases include **hydrogen** and **helium**.

Planetesimals

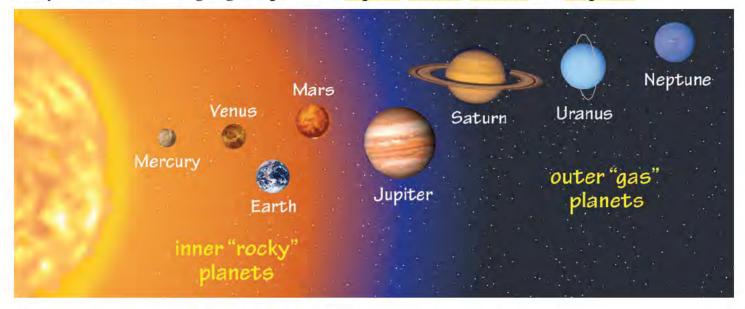
Within a **nebula**, dust particles collide and stick together to form larger particles. This process is called **accretion**. The particles grow from pebbles, to rocks, to boulders. The larger particles are called **planetesimals**.





Rocky & Gaseous Planet Formation

Rocky planetesimals near the Sun grew by accretion to become the rocky planets—Mercury, Venus, Earth and Mars. In the colder outer parts of the solar system, frozen gases accreted to rocky cores to form the gas giant planets—Jupiter, Saturn, Uranus and Neptune.

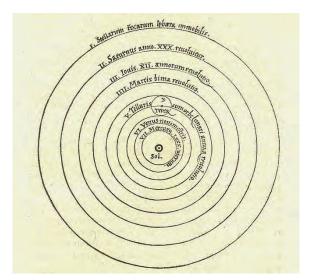




GEOCENTRIC & HELIOCENTRIC SYSTEMS

Geocentric Universe Concept

Before the invention of the **telescope**, the solar system was described in relation to the Earth. In **140 AD**, the Greek astronomer **Ptolemy** believed that the Earth was the center of the universe, and the Sun and planets revolved around the Earth. This is the **geocentric concept** of the universe.



Copernican/heliocentric model



Ptolemaic/geocentric model

Heliocentric Universe Concept

In **1543** Copernicus published a theory about the universe. He claimed that the Sun is the center of the universe and all the known planets revolved around the Sun. This is the heliocentric concept of the universe.

A Heliocentric Solar System

Today we know that the Sun is the center of our **solar system** – a collection of planets orbiting a single star, the **Sun**.

